

Supplemental materials: Competing Orders in a Nearly Antiferromagnetic Metal

I. DETAILS ON THE MONTE CARLO SIMULATIONS

A. Determinantal quantum Monte Carlo setup

The action Eq. (1) of the main text defines the partition function

$$Z = \int D(\vec{\varphi}, \bar{\psi}, \psi) e^{-S_\varphi - S_F} = \int D\vec{\varphi} e^{-S_\varphi} \text{Tr}_\psi [e^{-S_F}], \quad (\text{S1})$$

which we now bring into a form amenable to standard determinantal quantum Monte Carlo (DQMC) methods [36, 37] as they are presented in several pedagogical texts [38, 39, 61, 62]. We also describe which measures need to be taken to attain a computational time complexity no worse than the optimal $O(\beta\mathcal{N}^3)$, where $\beta = 1/T$ is the inverse temperature and $\mathcal{N} = L^2$ the number of lattice sites.

To allow for an efficient numerical evaluation of the trace in fermionic Fock space remaining in Eq. (S1), we discretize imaginary time $\tau = \ell\Delta\tau$, $\beta = m\Delta\tau$ ($\Delta\tau = 0.1$), and after a symmetric Suzuki-Trotter decomposition we obtain

$$Z = \int D\vec{\varphi} e^{-\Delta\tau \sum_{\ell=1}^m L_\varphi(\ell\Delta\tau)} \text{Tr}_\psi \left[\prod_{\ell=1}^m \hat{B}_\ell \right] + O(\Delta\tau^2). \quad (\text{S2})$$

Here the operators \hat{B}_ℓ are given by

$$\hat{B}_\ell = e^{-\frac{1}{2}\Delta\tau\psi^\dagger K\psi} e^{-\Delta\tau\psi^\dagger V_\ell\psi} e^{-\frac{1}{2}\Delta\tau\psi^\dagger K\psi}, \quad (\text{S3})$$

with non-commuting matrices K and V_ℓ and vectors of fermionic operators

$$\psi^\dagger = (\psi_{\alpha i \sigma}^\dagger) = (\psi_{x1\uparrow}^\dagger, \dots, \psi_{x\mathcal{N}\uparrow}^\dagger, \psi_{y1\downarrow}^\dagger, \dots, \psi_{y\mathcal{N}\downarrow}^\dagger, \quad (\text{S4}) \\ \psi_{x1\downarrow}^\dagger, \dots, \psi_{x\mathcal{N}\downarrow}^\dagger, \psi_{y1\uparrow}^\dagger, \dots, \psi_{y\mathcal{N}\uparrow}^\dagger).$$

Explicitly, K and V_ℓ are given by

$$K_{ij, \alpha\alpha', ss'} = \delta_{ss'} \delta_{\alpha\alpha'} (-t_{\alpha, s, ij} - \mu \delta_{ij}), \\ V_{\ell; ij, \alpha\alpha', ss'} = \lambda [\sigma_1]_{\alpha\alpha'} \delta_{ij} [\vec{s} \cdot \vec{\varphi}_i(\ell)]_{ss'}. \quad (\text{S5})$$

In this equation the Pauli matrix σ_1 acts on flavor indices, while the Pauli matrices \vec{s} act on spin indices. We allow the hopping constants t to depend on spin in order to implement a generalized magnetic field as described in Sec. IB below. In this $O(2)$ -symmetric model we have $\vec{\varphi} = (\varphi^1, \varphi^2)$. Carrying out the trace in Eq. (S2) yields

$$\text{Tr}_\psi \left[\prod_{\ell=1}^m \hat{B}_\ell \right] = \det \left[\mathbb{1} + \prod_{\ell=1}^m B_\ell \right] = \det G_\varphi^{-1} \quad (\text{S6})$$

with $B_\ell = e^{-\frac{1}{2}\Delta\tau K} e^{-\Delta\tau V_\ell} e^{-\frac{1}{2}\Delta\tau K}$ [39]. The matrix G_φ is the equal-time Green's function evaluated for one bosonic spin configuration $\{\vec{\varphi}_i(\ell)\}$. After partitioning the matrix exponentials into $\mathcal{N} \times \mathcal{N}$ -sized blocks, they read

$$e^{-\frac{\Delta\tau}{2}K} = \text{diag} \left(e^{-\frac{\Delta\tau}{2}K_x^\uparrow}, e^{-\frac{\Delta\tau}{2}K_y^\downarrow}, e^{-\frac{\Delta\tau}{2}K_x^\downarrow}, e^{-\frac{\Delta\tau}{2}K_y^\uparrow} \right), \\ e^{-\Delta\tau V(\ell)} = \begin{pmatrix} C & S \\ S^* & C \\ & C & S^* \\ & S & C \end{pmatrix} = \begin{pmatrix} \tilde{V}(\ell) & \\ & \tilde{V}(\ell)^* \end{pmatrix} \quad (\text{S7})$$

with submatrices

$$C_{ij} = \delta_{ij} \cosh(\Delta\tau |\vec{\varphi}_j(\ell)|), \quad (\text{S8}) \\ S_{ij} = \delta_{ij} [i\varphi_j^2(\ell) - \varphi_j^1(\ell)] \sinh(\Delta\tau |\vec{\varphi}_j(\ell)|) / |\vec{\varphi}_j(\ell)|.$$

Under the condition

$$K_x^\uparrow = K_x^{\downarrow,*} \quad \text{and} \quad K_y^\downarrow = K_y^{\uparrow,*} \quad (\text{S9})$$

the Green's function decomposes into two blocks of size $2\mathcal{N} \times 2\mathcal{N}$:

$$G_\varphi = \begin{pmatrix} \tilde{G}_\varphi & \\ & \tilde{G}_\varphi^* \end{pmatrix}. \quad (\text{S10})$$

Hence we can write the partition function as

$$Z = \int D\vec{\varphi} e^{-\Delta\tau \sum_{\ell=1}^m L_\varphi(\ell\Delta\tau)} \left| \det \tilde{G}_\varphi^{-1} \right|^2 + O(\Delta\tau^2), \quad (\text{S11})$$

which now is in a form that can be evaluated by Monte Carlo sampling over space-time configurations $\{\vec{\varphi}_i(\ell)\}$. Note that the probability measure under the field integral is positive definite, which allows for efficient sign-problem-free Monte Carlo simulations. The $O(2)$ -symmetry allows us to restrict all fermionic evaluations to the $(x\uparrow, y\downarrow)$ -sector, which speeds up the most expensive computations by a factor of 8 in comparison to the $O(3)$ -model. From the matrix G_φ we can compute arbitrary fermionic equal-time observables via Wick's theorem and also access imaginary-time-displaced correlation functions after the application of matrices B_ℓ and B_ℓ^{-1} .

Generally, in the DQMC algorithm we frequently need to compute products of the matrices B_ℓ . While the exponentials of V_ℓ are sparse matrices and consequently their multiplication has a computational cost of $O(\mathcal{N}^2)$ only, even for electron hopping restricted to nearest-neighbor sites, the exponentials of the kinetic matrices K_α^s are densely filled, which raises the cost of a single multiplication to $O(\mathcal{N}^3)$. We avoid paying this cost by performing a ‘‘checkerboard’’ decomposition [61], where we divide the whole set of lattice bonds into two groups, so that $K_\alpha^{s(1,2)}$ are sums of commuting four-site hopping matrices and $K_\alpha^s = K_\alpha^{s(1)} + K_\alpha^{s(2)}$. Applying this decomposition for all α and s , we find

$$B_\ell = e^{-\Delta\tau K/2} e^{-\Delta\tau V_\ell} e^{-\Delta\tau K/2} \quad (\text{S12}) \\ \approx e^{-\Delta\tau K^{(1)}/2} e^{-\Delta\tau K^{(2)}/2} e^{-\Delta\tau V_\ell} e^{-\Delta\tau K^{(2)}/2} e^{-\Delta\tau K^{(1)}/2}$$

and do not introduce any error of higher order than that already present from the Suzuki-Trotter decomposition, yet save one power of \mathcal{N} in computational effort.

B. Controlling finite-size effects

Simulations of metallic systems at low temperatures are particularly susceptible to strong finite-size effects. Since our numerical methods limit us to the study of finite lattices, reducing the severity of these effects is very important. A dramatic reduction of finite-size effects can be obtained in the presence of a perpendicular magnetic field [40]. In our simulations we thread a single magnetic flux quantum Φ_0 through the system, making sure not to break condition (S9) in order not to re-introduce a sign problem. Specifically, we add Peierls phase factors to the hopping terms of the kinetic operator:

$$t_{\alpha,s,ij}\psi_{x,i,s}^\dagger\psi_{x,j,s} \rightarrow e^{iA_{ij}^{\alpha s}} t_{\alpha,s,ij}\psi_{\alpha,i,s}^\dagger\psi_{\alpha,j,s} \quad \text{with} \quad A_{ij}^{\alpha s} = \frac{2\pi}{\Phi_0} \int_{\mathbf{r}_i}^{\mathbf{r}_j} d\mathbf{x} \cdot \mathbf{A}^{\alpha s} \quad (\text{S13})$$

and choose the Landau gauge $\mathbf{A}^{\alpha s}(\mathbf{x}) = -B^{\alpha s}x_2\hat{\mathbf{e}}_1$. The sign of the magnetic field depends on flavor and spin indices α, s and its magnitude is the smallest possible on the periodic $L \times L$ lattice:

$$B^{x\uparrow} = B^{y\downarrow} = -B^{x\downarrow} = -B^{y\uparrow} = \frac{\Phi_0}{L^2}. \quad (\text{S14})$$

Note that as $L \rightarrow \infty$ the original hopping constants are restored. To maintain translational invariance in presence of the magnetic flux we impose special boundary conditions in the $\hat{\mathbf{e}}_2$ -direction

$$\psi_{\alpha,\mathbf{r}+L\hat{\mathbf{e}}_2,s} = \psi_{\alpha,\mathbf{r},s} \exp\left(\frac{2\pi i}{\Phi_0} B^{\alpha s} L r_1\right), \quad (\text{S15})$$

while we retain regular periodic boundary conditions in $\hat{\mathbf{e}}_1$ -direction. Explicitly, for nearest-neighbor hopping, the phases read

$$A_{ij} = \begin{cases} -\frac{2\pi}{\Phi_0} B^{\alpha s} i_2 & \text{if } i_1 = 0, \dots, L-2 \text{ and } j_1 = i_1 + 1 \\ & \text{or } i_1 = L-1 \text{ and } j_1 = 0, \\ +\frac{2\pi}{\Phi_0} B^{\alpha s} i_2 & \text{if } i_1 = 1, \dots, L-1 \text{ and } j_1 = i_1 - 1 \\ & \text{or } i_1 = 0 \text{ and } j_1 = L-1, \\ +\frac{2\pi}{\Phi_0} B^{\alpha s} L i_1 & \text{if } i_2 = L-1 \text{ and } j_2 = 0, \\ -\frac{2\pi}{\Phi_0} B^{\alpha s} L i_1 & \text{if } i_2 = 0 \text{ and } j_2 = L-1, \\ 0 & \text{otherwise,} \end{cases} \quad (\text{S16})$$

where the lattice site vectors are $\mathbf{r}_i = (i_1, i_2)$ and $\mathbf{r}_j = (j_1, j_2)$, which we index from 0 to $L-1$ in each direction.

C. Local and global updates

The foundation of our Monte Carlo simulations of the lattice field theory (S11) is the Metropolis algorithm, where a proposed change of a bosonic field configuration $\{\vec{\varphi}\} \rightarrow$

$\{\vec{\varphi}'\}$ is accepted with probability

$$p = \min \left\{ 1, e^{-(S'_{\varphi} - S_{\varphi})} \left| \frac{\det \tilde{G}_{\varphi}}{\det \tilde{G}'_{\varphi}} \right|^2 \right\}. \quad (\text{S17})$$

The most important type of proposed changes is the local update. For one sweep of local updates we attempt to change the orientations and lengths of individual bosonic spins $\vec{\varphi}_i(\ell)$ chosen sequentially from the space-time lattice. During an initial equilibration phase we generally tune the size of the box from which the new spin is chosen such that about 50% of all local updates are accepted. For local updates the determinant ratio in (S17) is given by

$$\det \tilde{G}_{\varphi} / \det \tilde{G}'_{\varphi} = \det[\mathbb{1} + \Delta(\mathbb{1} - \tilde{G}_{\varphi}(\ell))], \quad (\text{S18})$$

where $\tilde{G}_{\varphi}(\ell) \equiv [\mathbb{1} + \tilde{B}_{\ell} \cdots \tilde{B}_1 \tilde{B}_m \cdots \tilde{B}_{\ell+1}]^{-1}$ and $\Delta \equiv \tilde{B}'_{\ell} \tilde{B}_{\ell}^{-1} - \mathbb{1} = e^{-\Delta\tau\tilde{V}'_{\ell}} e^{\Delta\tau\tilde{V}_{\ell}} - \mathbb{1}$. We find that an expansion by minors reduces the determinant in Eq. (S18) to that of a 2×2 -matrix $M = \mathbb{1}_2 + (\mathbb{1}_2 - \tilde{G}^i) \cdot \Delta^i$, where $\Delta^i = \Delta[i :: \mathcal{N}, i :: \mathcal{N}]$, $\tilde{G}^i = \tilde{G}_{\varphi}(\ell)[i :: \mathcal{N}, i :: \mathcal{N}]$, and the slice index notation corresponds to the four sole non-zero entries of Δ . Thus the acceptance probability can be computed in constant time.

After an accepted local update the Green's function matrix must be updated:

$$\tilde{G}'_{\varphi}(\ell) = \tilde{G}_{\varphi}(\ell)[\mathbb{1} + \Delta(\mathbb{1} - \tilde{G}_{\varphi}(\ell))]^{-1}, \quad (\text{S19})$$

where we can again exploit the sparseness of Δ and replace the inversion of the $2\mathcal{N} \times 2\mathcal{N}$ -matrix by that of a 2×2 -matrix if we make use of the Sherman-Morrison-Woodbury formula [63]. In this manner we find

$$\tilde{G}'_{\varphi}(\ell) = \tilde{G}_{\varphi}(\ell) + (\tilde{G}_{\varphi}(\ell)[:, i :: \mathcal{N}] \cdot \Delta^i) \cdot (M^{-1} \cdot \{(G_{\varphi}(\ell) - \mathbb{1})[i :: \mathcal{N}, :]\}), \quad (\text{S20})$$

which, if the matrix products are carried out in the order indicated by the parentheses, can be evaluated with only $O(\mathcal{N}^2)$ operations. In consequence, one total sweep of local updates has a time complexity of $O(\beta\mathcal{N}^3)$.

Using only these local updates, Monte Carlo simulations of this model require relatively long thermalization periods without measurements to equilibrate the system and then generally show long statistical autocorrelation times, which are amplified near the magnetic phase transition by critical slowing down. To counteract these effects we adopt two tools: a simple global update and a replica exchange mechanism.

The global update move consists of adding a constant random displacement $\vec{\delta}$ to all bosonic spins: $\vec{\varphi}_i(\ell) \rightarrow \vec{\varphi}_i(\ell) + \vec{\delta}$. To evaluate the acceptance probability (S17) we compute \tilde{G}'_{φ} and its determinant from scratch, which takes $O(\beta\mathcal{N}^3)$ operations. At times we found it also helpful to combine this move with the Wolff single cluster algorithm [64]. Here we ignore the fermionic part of the action while we construct and flip a cluster of spins, then we add the global displacement, and finally we decide on accepting the joint move according to Eq. (S17).

D. Replica exchange

For the replica exchange or parallel tempering scheme [65, 66] we consider an extended ensemble composed of multiple grand-canonical ensembles with the parameter r in S_φ taking on different values $r_1 < r_2 < \dots < r_K$ such that the partition function is given by a product $\mathcal{Z} = \prod_{\kappa=1}^K Z(r_\kappa)$,

$$\mathcal{Z} = \int D(\vec{\varphi}_1, \dots, \vec{\varphi}_K) \prod_{\kappa=1}^K e^{-S_\varphi(r_\kappa, \{\vec{\varphi}_\kappa\})} \left| \det \tilde{G}_{\varphi_\kappa}^{-1} \right|^2, \quad (\text{S21})$$

where $\tilde{G}_{\varphi_\kappa}$ does not depend on r_κ . In the Monte Carlo simulation we then have in parallel a separate replica of the system for every r_κ , each being represented by a different system configuration. The control parameter r is treated as a dynamical variable by allowing exchanges of the configurations between replicas with different parameter values. In this way shorter autocorrelation times at high r can be utilized to accelerate the simulation across the phase transition and in the low- r region. To achieve this we need to construct a Monte Carlo move between replicas, which will supplement the single-replica local and global updates that are still carried out as in regular canonical simulations. In such an update we propose the exchange of configurations $\{\vec{\varphi}\}$ and $\{\vec{\varphi}'\}$ between the κ -th and η -th replicas. To ensure detailed balance we require

$$P(\dots, \vec{\varphi}, r_\kappa, \dots, \vec{\varphi}', r_\eta, \dots) W(\vec{\varphi}, r_\kappa | \vec{\varphi}', r_\eta) = P(\dots, \vec{\varphi}', r_\kappa, \dots, \vec{\varphi}, r_\eta, \dots) W(\vec{\varphi}', r_\kappa | \vec{\varphi}, r_\eta), \quad (\text{S22})$$

where $P(\vec{\varphi}_1, r_1, \dots, \vec{\varphi}_K, r_K)$ is the equilibrium probability of a set of system configurations $\{\vec{\varphi}_\kappa\}$ associated to parameters r_κ in the extended ensemble and W is the transition probability for a replica configuration exchange. The ratio of these transition probabilities is

$$\frac{W(\vec{\varphi}, r_\kappa | \vec{\varphi}', r_\eta)}{W(\vec{\varphi}', r_\kappa | \vec{\varphi}, r_\eta)} = \frac{e^{-S_\varphi(r_\kappa, \{\vec{\varphi}'\}) - S_\varphi(r_\eta, \{\vec{\varphi}\})}}{e^{-S_\varphi(r_\kappa, \{\vec{\varphi}\}) - S_\varphi(r_\eta, \{\vec{\varphi}'\})}} = e^{-\Delta}, \quad (\text{S23})$$

where $\Delta = (r_\kappa - r_\eta) \cdot \frac{\Delta\tau}{2} \sum_{\ell,i} [\vec{\varphi}'_i(\ell)^2 - \vec{\varphi}_i(\ell)^2]$ and we note that the fermion determinants have canceled. To fulfill the relation (S23) we choose exchange probabilities according to the Metropolis criterion

$$W(\vec{\varphi}, r_\kappa | \vec{\varphi}', r_\eta) = \min \{1, e^{-\Delta}\}. \quad (\text{S24})$$

In our simulations we only propose exchanges between adjacent pairs of control parameter values. At high temperatures we achieve good diffusion with a simple linear spacing of the values of r . At lower temperatures, however, the magnetic phase transition constitutes a more significant barrier to the random walk in r -space. Here we have used a feedback-optimized distribution of r -values [67, 68], which effectively clusters the r_κ around $r_{\text{SDW}}(T)$, easing diffusion and significantly lowering autocorrelation times. Since the exchange algorithm following Eq. (S23) does not require the recomputation of Green's functions or the evaluation of their determinants, it poses very little overhead in computation or communication. This allows us to perform a replica-exchange sweep

after every single sweep of canonical updates, which has been very beneficial for obtaining sufficient statistics to resolve the magnetic phase diagram.

E. Time series reweighting

The structure of the action, Eq. (1) of the main text, where the r -dependence is fully contained in the bosonic part S_φ , allows to easily relate the canonical probability distribution of a configuration $\{\vec{\varphi}\}$ at a tuning parameter value r , $p_r(\vec{\varphi})$, to the distribution at another value r' : $p_{r'}(\vec{\varphi}) \propto e^{-(r'-r)E(\vec{\varphi})} p_r(\vec{\varphi})$, where $E(\vec{\varphi}) = \frac{\Delta\tau}{2} \sum_{\ell,i} \vec{\varphi}_i(\ell)^2$. From this relation one finds an expression for the expectation value of an observable \mathcal{O} at r' in terms of expectation values at r , which in turn can be estimated by time series averages from a Monte Carlo simulation carried out at r :

$$\langle \mathcal{O} \rangle_{r'} = \frac{\langle \mathcal{O} e^{-(r'-r)E} \rangle_r}{\langle e^{-(r'-r)E} \rangle_r} \approx \frac{\sum_n \mathcal{O}_n e^{-(r'-r)E_n}}{\sum_n e^{-(r'-r)E_n}}, \quad (\text{S25})$$

where n goes over the series of measured samples and \mathcal{O}_n and E_n are computed from the same system configuration. This reweighting procedure [69] is effective over quite a wide range around r .

From our replica exchange simulations we have Monte Carlo data for multiple close values of r . Building on the observation (S25), we can use the combined information from these time series for r_κ , $\kappa = 1, \dots, K$, to obtain improved observable estimates at $r_1 \leq r \leq r_K$ by multiple histogram reweighting [70, 71]. To do so we write the expectation value as

$$\langle \mathcal{O} \rangle_r = \frac{\int dE \Omega(E) e^{-rE} \mathcal{O}(E)}{\int dE \Omega(E) e^{-rE}} \quad \text{with} \\ \mathcal{O}(E) = \frac{\int D\vec{\varphi} \delta[E(\vec{\varphi}) - E] \mathcal{O}[\vec{\varphi}]}{\int D\vec{\varphi} \delta[E(\vec{\varphi}) - E]}, \quad (\text{S26})$$

where all non- r -dependent parts of the action are contained in the density of states $\Omega(E)$. We discretize E into levels E_α spaced ΔE apart and search the optimal estimator for $\Omega(E_\alpha)$, which reads

$$\hat{\Omega}_\alpha = \frac{\sum_\kappa H_{\alpha\kappa} [g_{\alpha\kappa} (1 - \Delta E \hat{\Omega}_\alpha e^{-r_\kappa E_\alpha + f_\kappa})]^{-1}}{\sum_\kappa M_\kappa \Delta E e^{-r_\kappa E_\alpha + f_\kappa} [g_{\alpha\kappa} (1 - \Delta E \hat{\Omega}_\alpha e^{-r_\kappa E_\alpha + f_\kappa})]^{-1}}. \quad (\text{S27})$$

Here $H_{\alpha\kappa}$ is the count of samples with $E \in [E_\alpha, E_\alpha + \Delta E)$ in the time series with $r = r_\kappa$, $g_{\alpha\kappa}$ is a statistical inefficiency factor related to the integrated autocorrelation time of the indicator function for this count, M_κ is the total number of samples for r_κ , and $f_\kappa = -\ln Z(r_\kappa)$ is given by

$$f_\kappa = -\ln \sum_\alpha \hat{\Omega}_\alpha \Delta E e^{-r_\kappa E_\alpha}. \quad (\text{S28})$$

Empirically, we find it adequate to set $g_{\alpha\kappa} \equiv 1$. Iteration of Eqs. (S27) and (S28) yields a converged estimate of $\hat{\Omega}_\alpha$ and

following Eq. (S26) we compute the estimate of $\langle O \rangle_r$ as a weighted average of the time series of \mathcal{O} for the different r_κ :

$$\hat{\mathcal{O}}(r) = \frac{\sum_{\kappa=1}^K \sum_{n=1}^{M_\kappa} \mathcal{O}_{\kappa n} w_{\kappa n}(r)}{\sum_{\kappa=1}^K \sum_{n=1}^{M_\kappa} w_{\kappa n}(r)} \quad (\text{S29})$$

with weights

$$w_{\kappa n}(r) = \sum_{\alpha} \frac{\psi_{\alpha \kappa n} \hat{\Omega}_{\alpha} e^{-r E_{\alpha}}}{\sum_{\kappa} H_{\alpha \kappa}}, \quad (\text{S30})$$

where $\psi_{\alpha \kappa n}$ is the indicator function for $E \in [E_{\alpha}, E_{\alpha} + \Delta E]$ evaluated at the n -th sample of the time series for r_κ .

The multiple histogram reweighting method allows us to finely interpolate between the original values r_κ of our simulations. In addition it provides a reduction of statistical error bars in the reweighted estimates compared to averages from single time series. In this work we have used the method for bosonic observables related to the magnetic transition, although it can be extended to all fermionic observables.

II. MAGNETIC TRANSITION

In the thermodynamic limit the model described by Eq.(1) of the main text cannot show magnetic long-range order at any $T > 0$ as stated by the Mermin-Wagner theorem [72]. Nevertheless, a finite-temperature phase transition of the Berezinsky-Kosterlitz-Thouless (BKT) type [73–75] is not precluded in this $O(2)$ -symmetric model. Defining a local magnetization density $\vec{m}_i = \frac{1}{\beta} \int_0^\beta d\tau \vec{\varphi}_i(\tau)$ and the total magnetization density $\vec{m} = \frac{1}{L^2} \sum_i \vec{m}_i$, in such a scenario \vec{m} vanishes in the thermodynamic limit $L^2 \rightarrow \infty$ even below the transition temperature T_{SDW} , where only finite systems will have quasi-long-range order with $\langle |\vec{m}| \rangle \neq 0$. At temperatures approaching T_{SDW} from above, the correlation length ξ diverges exponentially

$$\xi \sim \exp(b(T - T_{\text{SDW}})^{-\nu}), \quad T \rightarrow T_{\text{SDW}}^+, \quad (\text{S31})$$

with a critical exponent $\nu = 1/2$ and it stays infinite for all $T \leq T_{\text{SDW}}$, so that the entire low-temperature phase is critical. Spatial correlation functions of the local magnetization fluctuations decay exponentially above T_{SDW} and with a power law below T_{SDW} :

$$\langle \vec{m}_i \cdot \vec{m}_{i+\mathbf{x}} \rangle \sim \begin{cases} e^{-|\mathbf{x}|/\xi}, & T > T_{\text{SDW}}, \\ |\mathbf{x}|^{-\eta(T)}, & T \leq T_{\text{SDW}}. \end{cases} \quad (\text{S32})$$

The critical exponent η depends on temperature with $\eta(T_{\text{SDW}}) = 1/4$. Following e.g. Refs. [76, 77], we study the spin-density wave susceptibility

$$\chi = \beta \sum_i \langle \vec{m}_i \cdot \vec{m}_0 \rangle = \int_0^\beta d\tau \sum_i \langle \vec{\varphi}_i(\tau) \vec{\varphi}_0(0) \rangle = \beta L^2 \langle \vec{m}^2 \rangle \quad (\text{S33})$$

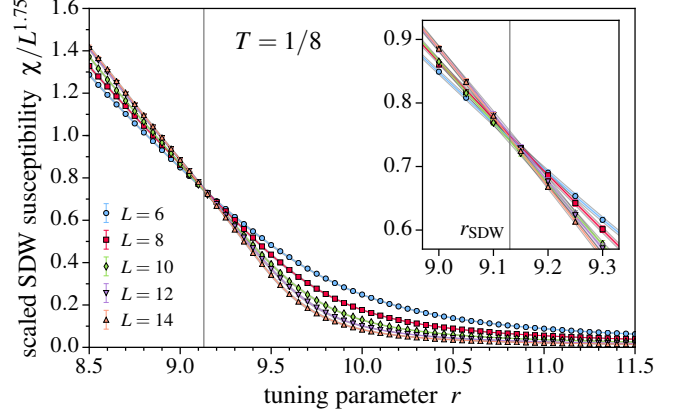


FIG. S1. Scaled SDW susceptibility $\chi/L^{2-\eta_c}$ with $\eta_c = 1/4$ for $T = 1/8$ and various system sizes. Symbols with error bars are estimates from single- r data. Continuous lines with surrounding error regions are results of the multiple-histogram reweighting analysis.

and from Eq. (S32) expect a finite-size scaling behavior like

$$\chi \sim L^{2-\eta} \quad (\text{S34})$$

with $\eta > 0$ for $T \leq T_{\text{SDW}}$ and slightly higher temperatures, where ξ still exceeds L . We identify points (r, T) in the phase diagram where Eq. (S34) can be fitted well to our data with $\eta \leq 1/4$ as belonging to the quasi-long-range ordered SDW phase.

At constant T we scan over r and fit the relation $\ln \chi = \alpha + (2 - \eta) \ln L$ to our data to determine $\eta(r)$. Then we search for $\eta(r_{\text{SDW}}) = 1/4$ to find where $T = T_{\text{SDW}}$. The reweighting technique described in Sec. IE provides us with high resolution in r to pinpoint r_{SDW} . In Fig. S1 we show that the intersection point of the scaled SDW susceptibility $\chi/L^{2-\eta_c}$ with $\eta_c = 1/4$ coincides approximately with this estimate for r_{SDW} . Fig. S2 illustrates the dependence of the estimated η on r , while Fig. S3 shows representative examples for fits with $\eta = 1/4$. As it is apparent there, the scaling relation (S34) fits our DQMC well for $T \gtrsim 1/16$, but for $T \leq 1/20$ we cannot find good agreement with the power law on the range of lattice sizes we have accessed. To account for a systematic error at these low temperatures we give a wider estimate of the error on r_{SDW} , allowing for values of $\eta \in [0, 0.5]$ (see Fig. S2b), while at higher temperatures we provide purely statistical error estimates computed from the variance-covariance matrix of the linear fit. A precise quantification of the systematic error in this finite-size scaling analysis would require system sizes L that are larger by orders of magnitude and hence out of computational reach. In Table I we summarize our results for $r_{\text{SDW}}(T)$ as determined from fits over five values $L = 6, \dots, 14$, which are also plotted in Fig. 1 in the main text, and show in comparison results for a reduced range $L = 8, \dots, 14$. In Fig. 1 the data points where we were not able to obtain a good fit ($T_{\text{SDW}} \leq 1/20$), are connected by bold lines.

The temperature below which the scaling law (S34) may be invalid lies under the superconducting T_c . There we have

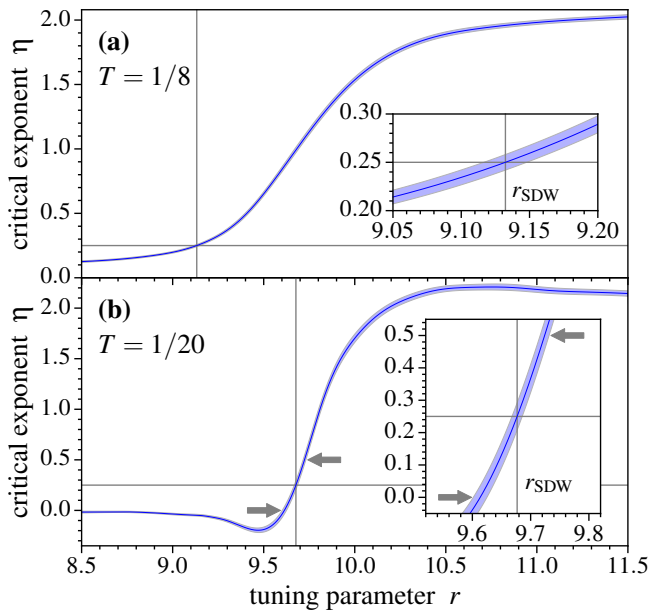


FIG. S2. Critical correlation exponent $\eta(r)$ at (a) $T = 1/8$ and (b) $T = 1/20$ for estimation of r_{SDW} , where $\eta = 1/4$. The continuous line shows the result of fitting $\ln \chi = \alpha + (2 - \eta) \ln L$ to the DQMC data, which has been interpolated by reweighting. The shaded surrounding region indicates the statistical error. The fits in (b) are of low quality. Here the arrows indicate a wider estimate of the error on r_{SDW} , allowing for $\eta \in [0, 0.5]$.

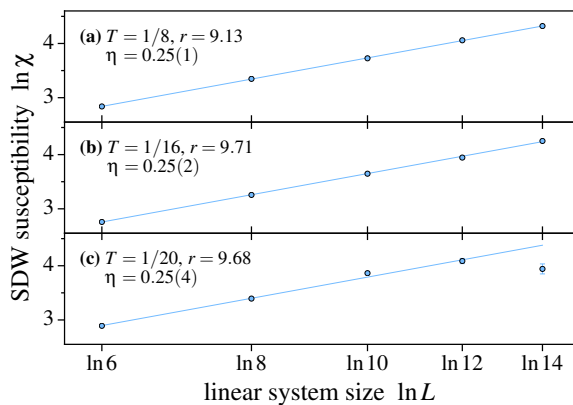


FIG. S3. Best fits of $\ln \chi = \alpha + (2 - \eta) \ln L$ with $\eta = 1/4$ for (a) $T = 1/8$, (b) $T = 1/16$, and (c) $T = 1/20$.

some indications that the magnetic transition could be weakly first-order and not of the BKT type. In extensive simulations at $T = 1/20$ for the largest system size $L = 14$ accessed by us the histograms of the finite-system magnetization density show a shallow double-peak structure when we tune r to an intermediate value between the magnetically quasi-long-range ordered and disordered phases, see Fig. S4. The location of this point is marked by a cross in Fig. 1 in the main text. If this dip grows deeper for larger systems, this bimodal distribution can be understood as a sign of phase coexistence at a first-order transition [78]. In our DQMC simulations close

TABLE I. Location of the SDW transition point r_{SDW} for different temperatures T as estimated by fitting $\ln \chi = \alpha + (2 - \eta) \ln L$ and searching for $\eta = 1/4$ for two ranges of system sizes $L = 6, \dots, 14$ ($n = 5$ data points) and $L = 8, \dots, 14$ ($n = 4$). $\chi_{\text{dof}}^2 = \chi^2 / (n - 2)$ is a measure to help with the estimation of the validity of the fit. For $T \leq 1/20$ the equation does not fit the data well and χ_{dof}^2 is larger than unity.

$1/T$	$L = 6, \dots, 14$		$L = 8, \dots, 14$	
	r_{SDW}	χ_{dof}^2	r_{SDW}	χ_{dof}^2
4	7.54(3)	0.6	7.6(1)	0.3
5	8.10(3)	1.4	8.07(5)	1.6
6	8.51(4)	1.3	8.499(2)	1.1
8	9.13(2)	0.9	9.12(3)	1.4
10	9.53(1)	0.4	9.52(3)	0.5
12	9.72(1)	1.8	9.73(3)	2.5
13	9.73(1)	0.1	9.73(1)	0.1
14	9.72(1)	4.0	9.76(1)	0.3
16	9.71(1)	0.5	9.71(1)	0.6
20	9.68(8)	10.2	9.7(1)	13.6
26	9.68(5)	11.0	9.66(7)	7.8
30	9.66(6)	4.4	9.62(9)	3.7

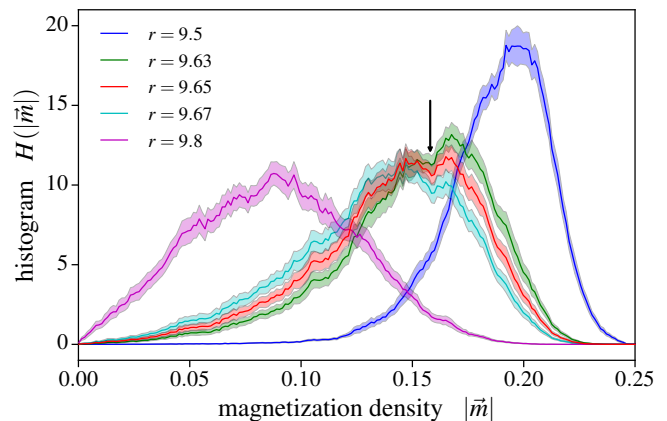


FIG. S4. Low-temperature histograms at $T = 1/20$ and $L = 14$ of the finite-system magnetization density $|\bar{m}|$ show a small suppression between two peaks at $r \approx 9.65$ close to the estimated location of the phase transition. This may be a signature of a weak first-order transition.

to the approximate transition point we also observe noticeably longer statistical autocorrelation times at $T \leq 1/20$ than at higher temperatures, which may be explained by the first-order transition and would also make it very cumbersome to obtain sufficient statistics to resolve these histograms for larger L .

III. DIAMAGNETIC RESPONSE AND THE IDENTIFICATION OF THE SUPERCONDUCTING T_c

In the appropriate gauge, the linear response of the system to a static, orbital magnetic field $B(\mathbf{q})$ is given by

$$j_x(\mathbf{q}) = -4K_{xx}(\mathbf{q})A_x(\mathbf{q}), \quad (\text{S35})$$

where $A_x(\mathbf{q}) = iB(\mathbf{q})/q_y$ is the vector potential in an appropriate gauge, and

$$K_{xx}(\mathbf{q}) \equiv \frac{1}{4} [\Lambda_{xx}(q_x \rightarrow 0, q_y = 0) - \Lambda_{xx}(\mathbf{q})]. \quad (\text{S36})$$

Here, Λ_{xx} is the current-current correlator

$$\Lambda_{xx}(\mathbf{q}) = \sum_i \int_0^\beta d\tau e^{-i\mathbf{q}\cdot\mathbf{r}_i} \langle j_x(\mathbf{r}_i, \tau) j_x(0, 0) \rangle, \quad (\text{S37})$$

and the current density operator is given by $j_x(\mathbf{r}_i) = \sum_{\alpha,s} it_{\alpha i s} \psi_{\alpha i s}^\dagger \psi_{\alpha j s} + \text{H.c.}$, where $\mathbf{r}_j = \mathbf{r}_i + \hat{x}$.

In the normal state, the magnetization is given by $-4\lim_{q_y \rightarrow 0} K_{xx}/q_y^2$. We note in passing that for general lattice models, the magnetic response can be of either sign. For the band parameters chosen in the text, the response in the non-interacting ($\lambda = 0$) case is paramagnetic.

To identify the superconducting transition, we employ the analysis of Ref. [42]. The superfluid density is given by [43]

$$\rho_s = \lim_{q_y \rightarrow 0} \lim_{L \rightarrow \infty} K_{xx}(q_x = 0, q_y) \quad (\text{S38})$$

Here, for convenience, we will use the notation $\rho_s(L) = K_{xx}(q_x = 0, q_y = 2\pi/L)$, whose limit when $L \rightarrow \infty$ is the superfluid density. At the BKT transition, the superfluid density changes discontinuously by a universal amount, $\Delta\rho_s = \frac{2T}{\pi}$. Figure S5 shows $\rho_s(L)$ across the phase diagram for multiple temperatures. For each temperature we identify the values of r at which $\rho_s(L) > \Delta\rho_s$ as the superconducting phase. The finite-size effects are not very substantial (except perhaps at large r at the lowest temperature $T = 0.025$), and are our main source of error in determining the superconducting phase boundary.

The analysis of the superfluid density does not rely on a particular ansatz for the superconducting order parameter. To determine the symmetry of the superconducting order parameter, we consider the uniform susceptibility $P_\eta(\mathbf{q} = 0)$, as defined in the main text. Close to the BKT transition, the susceptibility of the appropriate pairing channel scales as $L^{2-\eta}$, where η varies continuously with temperature, reaching the value $\eta = 0.25$ at T_c .

At low temperatures $P_-(\mathbf{q} = 0)$ is strongly dependent on L (see Fig. S6 and Fig. 2 in the main text). In contrast, $P_+(\mathbf{q} = 0)$ remains size-independent (shown in Fig. S7). Note also that the s -wave susceptibility is smaller by more than two orders of magnitude than the d -wave one. While we have not attempted to extract the transition temperature from the finite size scaling behavior of $P_-(\mathbf{q} = 0)$, it is clear that the pairing instability occurs in the d -wave channel.

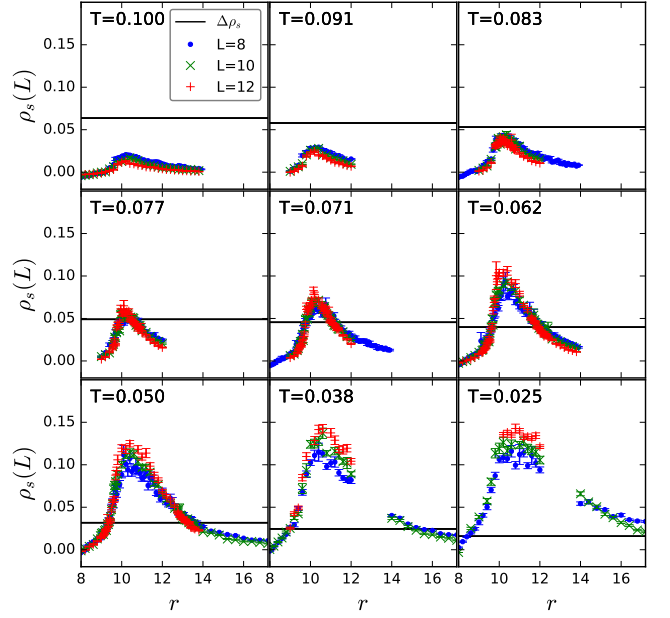


FIG. S5. $\rho_s(L)$, as defined in the text, for system sizes $L = 8, 10, 12$ across the phase diagram. The solid line indicates the universal value $\Delta\rho_s = \frac{2T}{\pi}$ expected at the BKT transition.

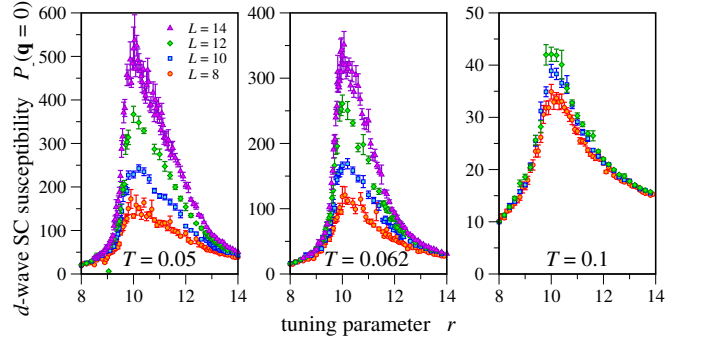


FIG. S6. d -wave superconducting susceptibility $P_-(\mathbf{q} = 0)$ as a function of r for different system sizes and temperatures.

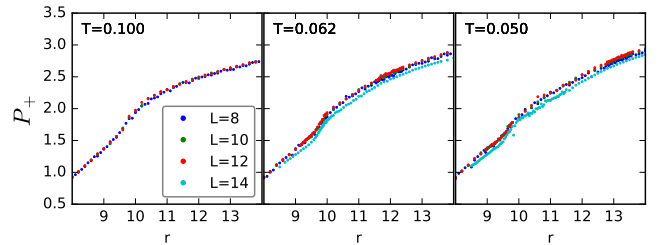


FIG. S7. Superconducting susceptibility in the s -wave channel, $P_+(\mathbf{q} = 0)$, across the phase diagram.

IV. CHARGE AND PAIR DENSITY WAVE SUSCEPTIBILITIES

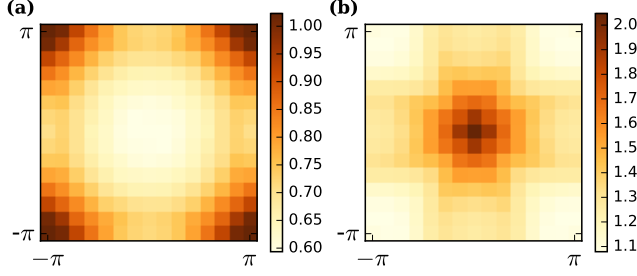


FIG. S8. (a) s -wave CDW and (b) s -wave PDW susceptibilities, as defined in the main text, across the Brillouin zone. Shown here is data for $L = 14$, $T = 0.083$, and $r = 10.4$.

In the main text we have focused on the d -wave CDW and PDW susceptibilities. The s -wave counterparts are shown in Fig. S8. Much like P_- , P_+ shows no structure at finite momenta. C_+ is peaked close to $\mathbf{q} = (\pi, \pi)$ (see also Fig. S9(a)), although the optimal \mathbf{q} can vary slightly with r (not shown). As the temperature is lowered, C_+ is at most moderately enhanced (see Fig. S9(b)), and its maximal value decreases with decreasing r .

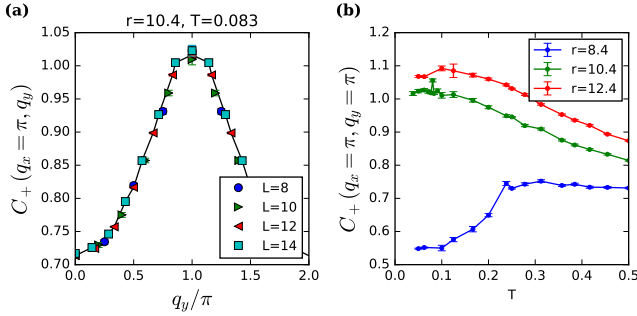


FIG. S9. (a) The s -wave CDW susceptibility versus momentum along the high-symmetry cut $\mathbf{q} = (\pi, q_y)$ for various system sizes. The solid line is a guide to the eye. (b) Temperature dependence of the CDW susceptibility at $\mathbf{q} = (\pi, \pi)$ for multiple values of r .

The quasi-one-dimensional character of the fermionic dispersion, shown in figure S10, can account for the enhancement of the CDW susceptibility. The CDW susceptibility for the non-interacting ($\lambda = 0$) case is shown in figure S11.

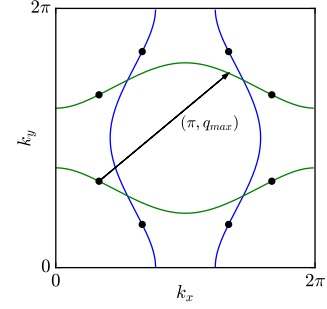


FIG. S10. The Fermi surface of the non-interacting ($\lambda = 0$) system. The hotspots of the SDW order are shown as black points. The arrow indicates the wavevector at which the CDW susceptibility is maximal.

Note that for this case $C_-(\mathbf{q}) = C_+(\mathbf{q})$. At low temperatures, $C_-(\mathbf{q})$ is peaked at $\mathbf{q} = (\pi, q_{\max}) = (\pi, 0.83\pi)$, similar to the interacting model. Note that this wavevector differs from any inter-hotspot wavevector. Due to the near-nesting of the Fermi surface (see figure S10), as the temperature is lowered, $C_-(\pi, q_{\max})$ increases and saturates at low temperatures, as shown in Fig. S11(b). Compared with C_- in the interacting case, shown in Fig. 6 of the main text, we see that the maximal CDW susceptibility in the interacting case is about 70% larger than the non-interacting one.

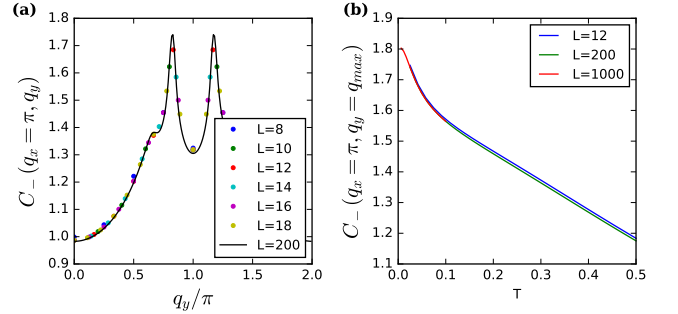


FIG. S11. (a) The non-interacting CDW susceptibility versus momentum along the high-symmetry cut $\mathbf{q} = (\pi, q_y)$ for various system sizes, shown here at $T = 0.025$ (b) Temperature dependence of the non-interacting CDW susceptibility.

To make sure the lack of tendency towards finite momentum charge order is not an artifact of the particular form-factor we have chosen, in figure S12 we examine the bond-density wave (BDW) susceptibilities $B_{\zeta, \eta}(\mathbf{q}) = \int d\tau \langle b_{\zeta, \eta}^\dagger(\mathbf{q}, \tau) b_{\zeta, \eta}(\mathbf{q}, 0) \rangle$. Here,

$$b_{\zeta, \eta}(\mathbf{q}) = \frac{1}{4} \sum_{s, \mathbf{k}} \left[\psi_{x, s, \mathbf{k}}^\dagger \psi_{x, s, \mathbf{k} + \mathbf{q}} + \eta \psi_{y, s, \mathbf{k}}^\dagger \psi_{y, s, \mathbf{k} + \mathbf{q}} \right] \left[\cos(k_x) + \cos(k_x + q_x) + \zeta \cos(k_y) + \zeta \cos(k_y + q_y) \right], \quad (\text{S39})$$

and η and ζ take the values ± 1 . The peaks in the BDW susceptibilities occur at momenta far from (π, π) , and show similar behavior to the non-interacting system (not shown). $B_{-, \zeta}$

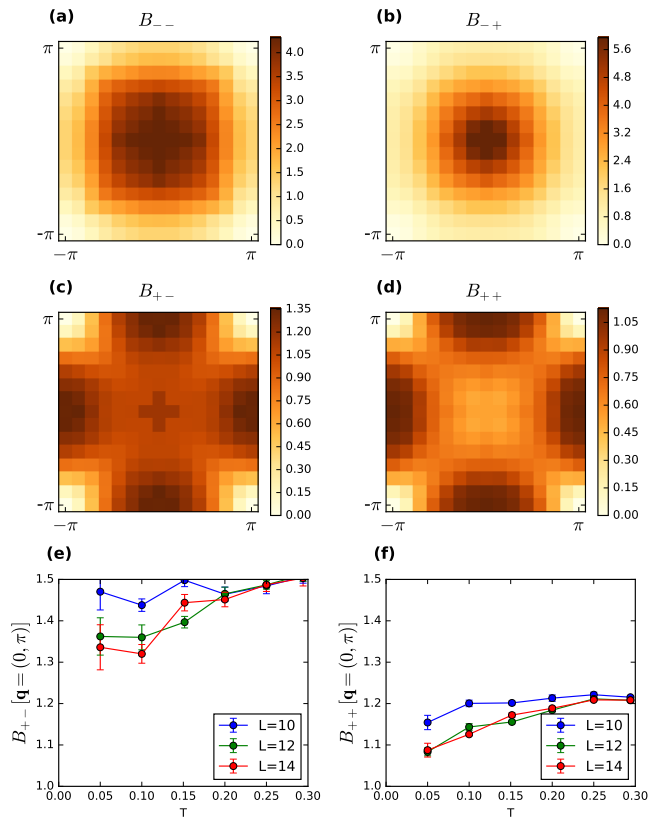


FIG. S12. (a-d) Bond-density wave susceptibilities across the Brillouin zone, shown here for $T = 0.1$, $L = 14$, $r = 10.4$. The value shown at $q = 0$ is interpolated from the four neighboring wave-vectors. (e-f) Temperature dependence of $B_{+, \eta}$ at the optimal wave-vector $\mathbf{q} = (0, \pi)$.

does not exhibit a peak at finite momentum. The temperature and finite-size dependence of $B_{+, \zeta}(\mathbf{q} = (0, \pi))$, shown in panels (e,f) of figure S12, are relatively weak.

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